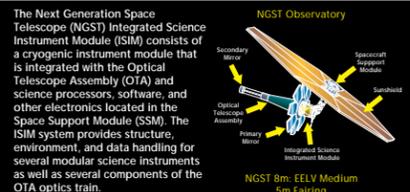




NEXT GENERATION NGST SPACE TELESCOPE



An ISIM baseline design and feasibility study is ongoing by GSFC. This pre-Phase A design was developed for integration with the Yardstick NGST architecture and packaging in a 5 m class EELV fairing. The goals of this study are to:

- [1] demonstrate mission science feasibility, [2] assess ISIM engineering and cost feasibility, [3] identify ISIM technology challenge areas, and [4] enable smart customer procurement of the NGST. In depth results from this work beyond those displayed here can be found at: <http://www701.gsfc.nasa.gov/isim/isim.htm>

The flight ISIM will be developed by a GSFC led Integrated Product Team (IPT) that will grow during Phase A/B to include members from the STScI, the NGST Prime Contractor, and science instrument development teams from European, Canadian, and NASA science communities. Science instruments will be competitively procured from the science community, and will be integrated into the ISIM by GSFC. The flight qualified ISIM will then be delivered by GSFC to the NGST Prime Contractor for observatory level integration.

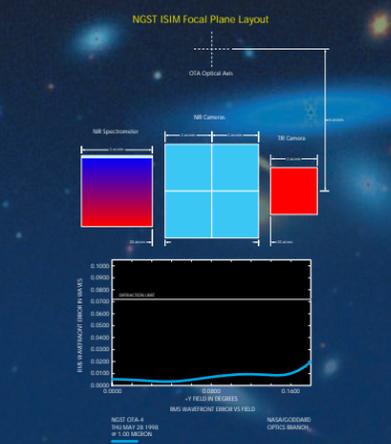
At the start of NGST Phase A (April, 1999), two competing prime contractors will begin development of separate NGST architectures, and the ISIM IPT will develop two ISIM designs corresponding to these architectures. Down selection to a single design will occur during April 01.

The ISIM team welcomes science community feedback. Contact the IPT lead: Matt Greenhouse, matt@stars.gsfc.nasa.gov.

ISIM Baseline Science Instruments

Instrument	Bandwidth	Bandwidth	FPA	Pixel Pitch	Pixel Scale	Aperture Control
Near-IR Camera (1 of 4)	0.8 - 5.0	4 x 2.5 Trued Beams	2048 x 4096	100	100	Quasi-beam splitter, Fov 2.2 x 2.2 arc-min
Near-IR Spectrometer	0.8 - 5.0	F = 300, 200 grating	2048 x 4096	21	100	Reflective IR beam splitter, 2048 pixel IR detector - 2 x 2 camera
Mid-IR Camera/Spectrometer	5 - 28	Local band filters	1024 x 1024	21	100	Micro-mirror array, 1024 pixel IR detector - 2 x 2 camera

1. A quasi-beam splitter (quipped mirror) approximates a 4 x 4 arc-min field of view (ovg) 4 identical cameras.

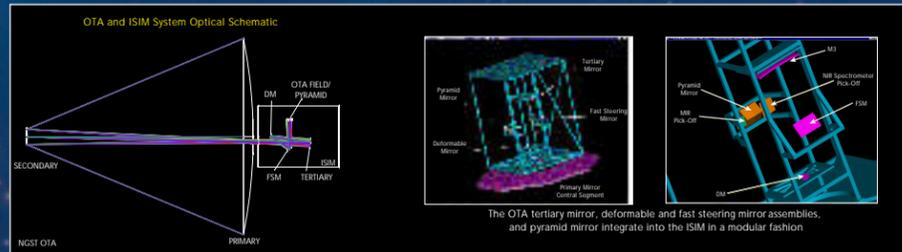


The ISIM instruments are located in an off-axis position. This configuration yields excellent image quality over a 24 arc-min diameter field.

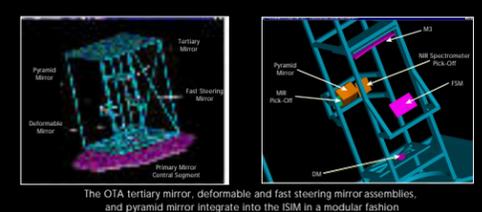
The NGST Yardstick Integrated Science Instrument Module (ISIM) Feasibility Study

M.A. Greenhouse, M. DiPirro, B. Federline, Jonathan P. Gardner, P. Guy, J. Hagopian, J. Hein, M. Jurotich, J. Lawrence, B. Martineau, J. Mather, E. Mentzell, S. Satyapal, D. Stanley, H. Teplitz (NASA GSFC), P. Bely, L. Petro, P. Stockman (STScI), R. Bug (Johns Hopkins U.), and R. Bitzel (Swales Aerospace)

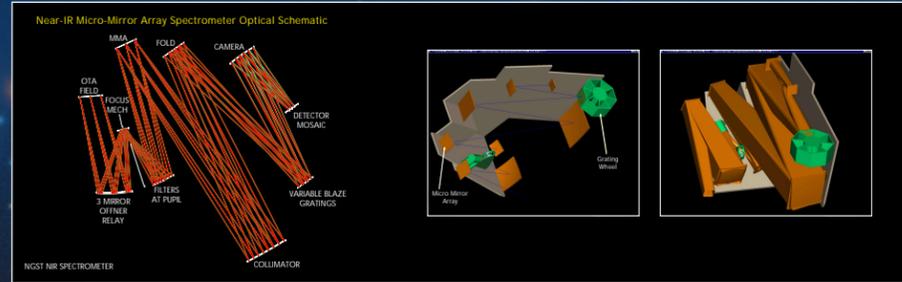
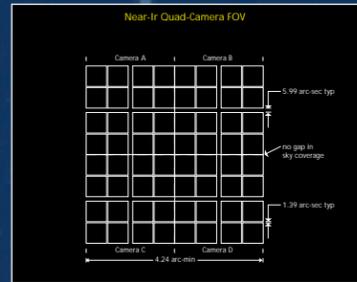
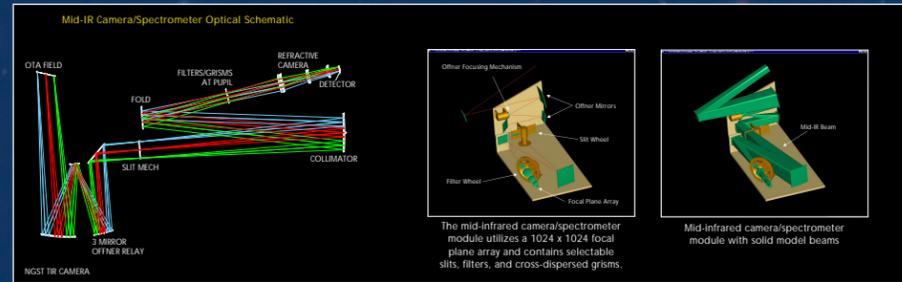
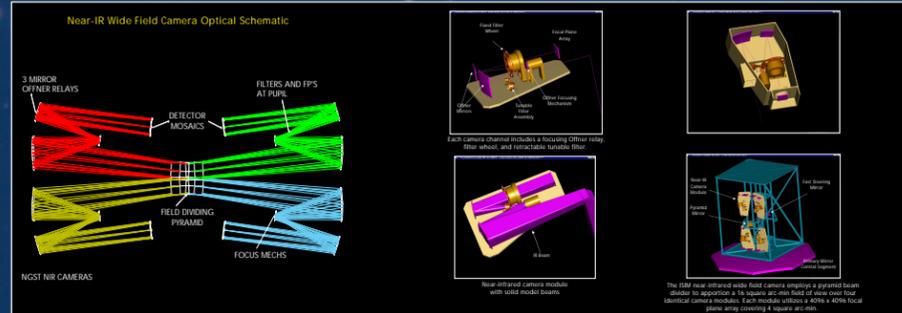
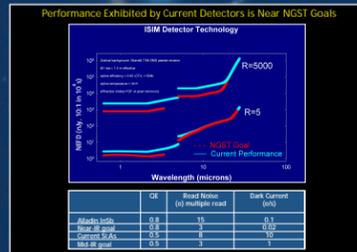
Optical Design



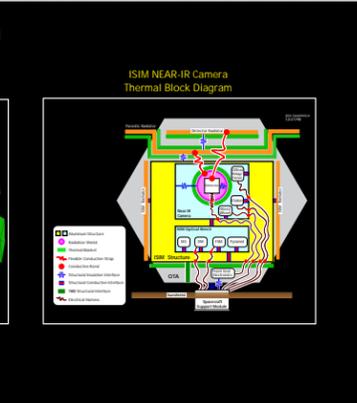
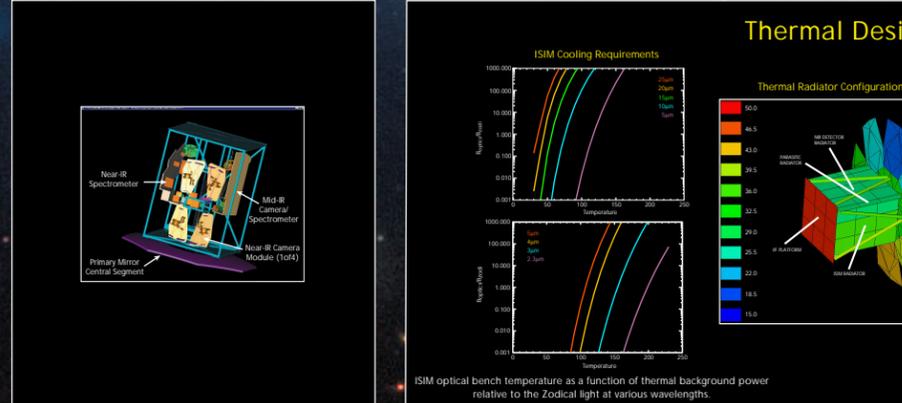
Opto Mechanical Layout



Focal Plane Arrays



- Detector Technology Development**
- Five technology development grants issued by NGST
 - Near-IR 0.8 - 5.0 um
 - Raytheon IRCoE & University of Rochester
 - 1024 x 1024 InSb, buttable to 4k mosaic, 30K operation
 - Rockwell Science Center & University of Hawaii
 - 2048 x 2048 HgCdTe and Si p-n diodes, buttable to 4k mosaic, 30K operation
 - Mid-IR
 - Raytheon IRCoE & NASA ARC & Cornell University
 - 512 x 512 SiAs, buttable to 1k mosaic, 6-8K operation, 5-20 um
 - Boeing Research & Technology Center
 - Si-Ga, 1k mosaic, 10-12K operation, 5-10 um
 - Rockwell Science Center & University of Rochester
 - HgCdTe, 25-30K operation, 5-10 um
 - Funded by Ball Aerospace
 - Raytheon IRCoE
 - 2048 x 2048 ROIC for InSb



ISIM optical bench temperature as a function of thermal background power relative to the Zodiical light at various wavelengths.